Improving the *Q* Factor of an Optical Atomic Clock Using Quantum Nondemolition Measurement

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(Received 24 July 2020; revised 20 September 2020; accepted 13 October 2020; published 15 December 2020)

Quantum nondemolition (QND) measurement is a remarkable tool for the manipulation of quantum systems. It allows specific information to be extracted while still preserving fragile quantum observables of the system. Here we apply cavity-based QND measurement to an optical lattice clock—a type of atomic clock with unrivaled frequency precision—preserving the quantum coherence of the atoms after readout with 80% fidelity. We apply this technique to stabilize the phase of an ultrastable laser to a coherent atomic state via a series of repeated QND measurements. We exploit the improved phase coherence of the ultrastable laser to interrogate a separate optical lattice clock, using a Ramsey spectroscopy time extended from 300 ms to 2 s. With this technique we maintain 95% contrast and observe a sevenfold increase in the clock's Q factor to 1.7×10^{15} .

DOI: 10.1103/PhysRevX.10.041052

Subject Areas: Atomic and Molecular Physics

I. INTRODUCTION

In quantum nondemolition (QND) measurement, an observable \hat{S} of a quantum system is coupled to an observable \hat{M} of a "meter" system, so that direct measurement of \hat{M} yields indirect information about \hat{S} . While the measurement of \hat{M} may perturb the state of the meter, the inferred value of the observable \hat{S} is conserved by the QND measurement [1]. QND measurements have given us a window on a wide range of quantum systems, including circuit quantum electrodynamics [2–4], solid-state spin qubits [5–7], mechanical oscillators [8,9], photons [10–13], nitrogen-vacancy centers [14], and trapped ions [15,16].

In this work we use QND measurement to observe cold Sr atoms in an optical lattice clock (OLC), in pursuit of metrological enhancements already demonstrated in Rband Cs-based magnetometers [17,18] and microwave atomic clocks [19–23]. Our work builds on recent demonstrations with Yb [24] and Sr [25,26] by applying QND measurement to a fully operational Sr OLC—an exceptionally stable and accurate type of clock [27–30] which is a prime candidate to underpin a future redefinition of the second in the International System of Units [31] as well as being a sensitive probe for geodesy [32,33] and physics beyond the standard model [34–37].

The OLC works by steering the frequency of an ultrastable laser, or "local oscillator" (LO), to match the frequency of the optical ${}^{1}S_{0}$ - ${}^{3}P_{0}$ clock transition in atomic Sr. The LO frequency is initialized close to resonance with the atomic clock transition, then a spectroscopy pulse is carried out on Sr atoms confined in an optical lattice in the ${}^{1}S_{0}$ ground state. At the end of the spectroscopy pulse, the frequency detuning between the LO and the atomic resonance is inferred by measuring the fraction of atoms excited into the ${}^{3}P_{0}$ state. In earlier realizations of the OLC [27–30] the excitation fraction is measured using fluorescence detection, which destroys the atomic sample. Stabilization of the LO therefore requires new atomic samples to be prepared, interrogated, and measured in a repeated cycle. By contrast, in this work the excitation fraction is measured using QND methods, allowing the atoms to be recycled for another spectroscopy pulse immediately after measurement.

We carry out QND measurement in an OLC by surrounding the Sr atoms with a high-finesse optical cavity at 461 nm, the wavelength of the strong ${}^{1}S_{0}{}^{-1}P_{1}$ transition. The same optical cavity also supports a magic-wavelength optical lattice trap [38]. The 461 nm intracavity photons serve as a QND meter of the number of ground state atoms, experiencing a measurable phase shift due to dispersion from the ${}^{1}S_{0}{}^{-1}P_{1}$ transition. In this work we demonstrate that, for short probe times, the QND measurement is weak and therefore preserves with high fidelity the coherence of atoms prepared in a superposition of ${}^{1}S_{0}$ and ${}^{3}P_{0}$. This nondestructive detection enables operation of the OLC in new, more stable configurations, such as the "atom phase lock" (APL), in which the phase of the LO is stabilized to the phase evolution of the atoms. Here we show that the APL significantly

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improves the coherence time of the LO laser. Deploying the phase-locked LO in a second, cointerrogated OLC enables us to extend the Ramsey dark time *T*, thereby reducing the Fourier-limited linewidth of the atomic signal $\Delta \nu = 1/(2T)$. This leads to an increased *Q* factor—i.e., an increased ratio $Q = \nu_0/\Delta\nu$ between the clock transition frequency ν_0 and the spectroscopic linewidth—enhancing a key figure of merit impacting the measurement precision of the clock.

II. QUANTUM NONDEMOLITION MEASUREMENT IN AN OPTICAL LATTICE CLOCK

To operate the OLC, fermionic strontium (⁸⁷Sr) is laser cooled and loaded into a magic-wavelength, onedimensional optical lattice at 813 nm. A Ramsey spectroscopy sequence then maps phase or frequency errors of the LO, in this case an ultrastable laser at 698 nm (see Supplemental Material [39]), onto a population imbalance between the electronic ground state $|g\rangle$ (5 s^2 ¹S₀, $M_F = \pm 5/2$) and the long-lived excited state $|e\rangle$ (5 s^5p ³ P_0 , $M'_F = \pm 3/2$). Adopting a pseudospin formulation, this population imbalance is encoded in the observable \hat{S}_z , the *z* component of the collective spin of the system \hat{S} . The collective spin components can be defined as

$$\hat{S}_x = \frac{1}{2}(\hat{S}_{ge} + \hat{S}_{eg}),$$
 (1)

$$\hat{S}_{y} = \frac{1}{2i} (\hat{S}_{eg} - \hat{S}_{ge}),$$
 (2)

$$\hat{S}_z = \frac{1}{2}(\hat{S}_{ee} - \hat{S}_{gg}),$$
 (3)

where the operators $\hat{S}_{ij} = \sum_{k=1}^{N} |i\rangle_k |j\rangle_k$ are summed over all atoms in the sample.

For a typical OLC, S_z is measured destructively in a twostep process [27–30]. First, a strong transition at 461 nm from the ground state to an auxiliary state $(5s5p \ ^1P_1)$ is used to measure S_{gg} via fluorescence detection. The fluorescence pulse heats the ground state atoms, causing them to escape from the lattice. Next, excited state atoms are optically pumped into the ground state and the fluorescence detection is repeated, giving a measurement



FIG. 1. Overview of the QND measurement scheme. (a) Sketch of the dual-wavelength in-vacuum cavity used to trap the atoms and to carry out the QND measurement. Atoms are trapped at the 813 nm intensity maxima represented in red, while they also interact with the nearest blue- and red-detuned cavity modes at 461 nm represented in blue and purple. (b) Simplified level scheme for Sr showing the 461 nm transition used for nondestructive detection and the 698 nm optical clock transition. (c) Diagram of the optical setup used for the QND measurement, and a sketch of the optical spectrum transmitted through the Mach-Zehnder interferometer (MZI). The six probe frequency components generated by the electro-optic modulator (EOM) chain are depicted in blue, interacting with the cavity modes in gray which surround the atomic transition in purple. The padlocks represent Pound-Drever-Hall loops used to stabilize the laser frequencies and the cavity lengths.

of S_{ee} . From these two measurements, S_z is calculated and the result is used to correct the LO frequency.

For the OLC in this work, we instead implement a QND measurement of the ground state population using the same optical cavity used to create the one-dimensional lattice trap. The cavity is coated to support optical modes surrounding the 461 nm transition from the ground to the auxiliary state. In the dispersive limit, where the detuning Δ of the cavity mode from the atomic transition is much larger than the cavity decay rate ($\kappa = 2\pi \times 330$ kHz), the atomic decay rate ($\Gamma = 2\pi \times 680$ kHz, the auxiliary state can be adiabatically eliminated. What remains is an effective coupling between the ground state population S_{gg} and the photon number in the cavity mode $\hat{c}^{\dagger}\hat{c}$, described by the Hamiltonian [40,41]:

$$\hat{H}_c = \hbar g^2 \hat{c}^\dagger \hat{c} \hat{S}_{aa} / \Delta. \tag{4}$$

This gives rise to an atom-induced frequency shift of the cavity resonance $\delta\nu = \langle \hat{H}_c/h \rangle / \langle \hat{c}^{\dagger} \hat{c} \rangle$. The basic principle of the QND measurement is to drive the cavity with a weak input field at 461 nm, so that the reflected output field carries information about $\delta\nu$, and therefore acts as a meter for the number of atoms. The phase of the reflected field is measured destructively as a beat signal on a photodetector, giving a signal proportional to S_{gg} . To obtain S_z , which is needed to estimate the LO frequency error, the ground and excited state populations are swapped via a π pulse T at 698 nm and a second QND measurement of S_{qg} is performed.

Further technical details of the QND measurement scheme [42] are outlined in Fig. 1. In order to provide first-order immunity to cavity length fluctuations [25,43,44], we probe the difference in the atomic-induced frequency shift between two adjacent longitudinal cavity modes centered in frequency around the atomic transition. The optical field used to probe

the two cavity modes is generated by sending the 461 nm laser through a Mach-Zehnder interferometer (MZI) amplitude modulator biased to zero throughput and driven at a frequency $\Omega/2 = 2.09$ GHz, matching the 4.18 GHz free spectral range of the cavity. Additional sidebands at $\Omega/2 \pm 125$ MHz are applied using the MZI modulator, generating strong frequency components which are directly reflected from the cavity input mirror. The strong directly reflected sidebands interfere with the cavity-coupled probe sidebands at $\pm \Omega/2$, generating a Pound-Drever-Hall [45] beat signal at 125 MHz proportional to the phase shift induced on the probe sidebands due to the atom-induced cavity shift $\delta \nu$.

III. WEAK QND MEASUREMENT AND ATOM COHERENCE PRESERVATION

To a good approximation the value of S_z is conserved after the QND measurement, but other properties of the atomic system can be significantly altered. For example, a fundamental measurement backaction is exerted by photon shot noise in the probe beam, which generates an increase in the uncertainty of S_y as we extract information about S_z , in compliance with the uncertainty principle $\Delta S_y \Delta S_z \ge$ $\langle |S_x| \rangle/2$. In practice, however, two other technical effects are much larger for the QND scheme in this work: (1) the photon scatter Γ_{sc} into free space and (2) the inhomogeneous ac Stark shift Δ_{ac} . Here, we discuss how these two forms of measurement backaction cause decay in the atom coherence S_x . We develop a model for the decoherence, and we present experimental data demonstrating weak QND readout of S_z while preserving S_x with 80% fidelity.

The scatter and the ac Stark shift depend on the radial position ρ and the position z along the cavity axis, according to

$$\Gamma_{\rm sc}(\rho,z) = \langle \Gamma_{\rm sc}(0,z) \rangle_z e^{-2\rho^2/w_0^2} \bigg[(\cos^2 kz + \sin^2 kz) + \frac{2\Delta_{\rm sum}\Delta_{\rm diff}}{\Delta_{\rm sum}^2 + \Delta_{\rm diff}^2} (\cos^2 kz - \sin^2 kz) \bigg],\tag{5}$$

$$\Delta_{\rm ac}(\rho,z) = \langle \Gamma_{\rm sc}(0,z) \rangle_z e^{-2\rho^2/w_0^2} \left[\frac{\Delta_{\rm diff}}{\Gamma} \left(\cos^2 kz - \sin^2 kz \right) + \frac{\Delta_{\rm sum}}{\Gamma} \left(\cos^2 kz + \sin^2 kz \right) \right] \left(1 - \frac{2\Delta_{\rm sum}^2}{\Delta_{\rm sum}^2 + \Delta_{\rm diff}^2} \right), \tag{6}$$

where $\langle \rangle_z$ indicates a spatial average along $z, w_0 = 75 \ \mu \text{m}$ is the waist of the cavity mode, $\Delta_{\text{diff}} = (\Delta_{\text{blue}} - \Delta_{\text{red}})/2 = 2\pi \times 2.09$ GHz is the average magnitude of the cavity mode detuning, $\Delta_{\text{sum}} = (\Delta_{\text{blue}} + \Delta_{\text{red}})/2 = -2\pi \times 173$ MHz is the asymmetry of the cavity mode detuning, $\Gamma = 2\pi \times 30$ MHz is the transition linewidth, and $k = 2\pi/\lambda$ is the wave number of the probe. In both equations we have explicitly written separate terms proportional to $\cos^2 kz$ and $\sin^2 kz$, created by the red- and blue-detuned probe sidebands respectively close to the center of the optical cavity. Ideally we would simplify the equations by choosing $\Delta_{\text{sum}} = 0$, but in practice a small offset is enforced by the technical constraint that the cavity length must be tuned to support a magic-wavelength 813 nm lattice to carry out high-*Q* spectroscopy on the optical clock transition. None-theless, we still operate with $\Delta_{\text{diff}} \gg \Delta_{\text{sum}}$, such that Eq. (5) yields an approximately uniform photon scatter rate along *z* while Eq. (6) yields an inhomogeneous ac Stark shift varying as $\cos 2kz$.

In order to model the effect of Δ_{ac} and Γ_{sc} on the collective atomic spin components S_i , we simulate a sample of a few thousand individual spins at different positions ρ , z

and propagate each spin using optical Bloch equations. The position ρ of each atom is selected from a Gaussian distribution with standard deviation $\sigma_{\rho} = 35 \ \mu m$ corresponding to a radial temperature of 5 μ K, which has been determined experimentally through sideband spectroscopy [46]. Since the radial trap frequency is only 120 Hz, we treat ρ as fixed throughout the QND measurement pulse, which has duration t < 0.5 ms. The position of each atom along z is randomly selected from one of 2000 sites of the 813 nm lattice trap, matching the experimentally measured width of the cloud. Along z, the trap frequency 63 kHz is comparable to or faster than 1/t, so we make the approximation that the mean z position of each atom is fixed to the center of the lattice site, but we average the scatter rate and ac Stark shift over a thermal waist $\sigma_z = 50$ nm corresponding to the 4 μ K measured axial temperature.

We investigate the QND probe backaction experimentally using the sequence depicted in Fig. 2(a). A sample of 6×10^3 atoms is first prepared in a coherent state with $\langle S_x \rangle = N/2$ using a resonant $\pi/2$ pulse from the clock laser. The QND probe is then applied for a variable amount of time t. After this, a second $\pi/2$ pulse is applied from the clock laser, the phase of which is stepped by 0° or 180° with respect to the first pulse in order to map S_x to $\pm S_z$. Finally, a destructive measurement is carried out of S_z , from which the value of S_x just before the second $\pi/2$ pulse can be inferred. To provide insensitivity to small systematic offsets in the S_{7} measurement, the estimate of S_x is based on the difference in measured S_{z} between the two phases 0°, 180° of the final clock pulse. As observed in the data "without spin echo" in Fig. 2(b), the inhomogeneous ac Stark shift Δ_{ac} results in near-total loss of coherence at QND probe time $t = 100 \ \mu s$. However, the rapid decoherence can be largely reversed using a spin-echo protocol. In the "with spin echo" sequence, an additional π pulse is inserted with phase 90° after the first QND probe, followed by a second QND probe. We observe that the decoherence from the ac Stark shift is strongly suppressed by the spin echo, with residual exponential decay of S_x with a time constant 317 μ s when using 125 fW of cavity-coupled QND probe light. Since the π pulse inverts the ground and excited population, the difference between the two QND probe signals in the spin-echo sequence provides a value for S_z . Therefore, a spin-echo QND probe sequence with a total probe time $t = 60 \ \mu$ s can act as a weak measurement of S_z , creating a signal to stabilize the clock LO while maintaining coherence with 80% fidelity.

IV. INCREASING THE *Q* FACTOR VIA AN ATOM PHASE LOCK

QND measurement in an OLC enables several novel applications that are otherwise impossible using conventional fluorescence readout techniques. Here we pursue one such application—the atom phase lock—in which the phase noise of the LO is tracked and corrected for via repeated weak measurement of the collective atomic spin. We characterize the performance of the APL to one OLC (Sr2 [47]) using synchronous interrogation of a second OLC (Sr1 [48,49]) which has highly correlated sensitivity to fluctuations in the LO frequency and phase (see Fig. 3 and Supplemental Material [39]).

After loading approximately 1×10^4 atoms into the optical lattice, the APL is implemented in Sr2 following the scheme depicted in Fig. 4(a), which was originally



FIG. 2. Coherence preservation after the QND measurement. (a) Timing sequence used to measure coherence after QND measurement, and sketches of the atomic state in the Bloch sphere representation at each step of the sequence. The projection of atoms into the ground or excited state due to scattering of QND probe photons is represented by the shrinking radius of the Bloch sphere compared to its original size N/2 (gray halo). The final Bloch sphere (top right) shows the case where the final $\pi/2$ -pulse phase is $\phi = 0^{\circ}$. (b) The measurement time summed over the two QND probes, and fit an exponential decay with time constant 317 μ s (blue dashed line). The model for coherence decay without spin echo (green dashed line, also displayed in the inset) is described in the main text.



FIG. 3. Experimental setup and Sr1-Sr2 correlations. (a) LO light is distributed to Sr1 and Sr2 along separate optical paths, the lengths of which are actively stabilized. When the APL is engaged, phase corrections from Sr2 are applied to an acousto-optic modulator (AOM) shared by both systems, thereby increasing the coherence time of the light sent to Sr1. (b) Allan deviation of the frequency ratio between Sr1 and Sr2, with the APL disengaged and the OLCs independently stabilized using synchronous 300 ms Rabi pulses. Both clocks experience the same LO frequency fluctuations, resulting in highly correlated frequency corrections ν_1 , ν_2 . However, there are residual sources of noise—for example, linear Zeeman shift fluctuations, which are suppressed using a less sensitive Zeeman transition. After minimizing noise (see Supplemental Material [39]), the frequency instability approaches the quadrature sum of QPN from both clocks, $4 \times 10^{-17}/\sqrt{\tau}$. The cycle time is 1.75 s and atom numbers are 7×10^3 and 1.3×10^4 atoms in Sr1 and Sr2, respectively.



FIG. 4. Enhanced Ramsey spectroscopy via APL. (a) Timing sequence for the synchronous spectroscopy scheme, and Bloch sphere representation of the atomic state during the APL sequence. Atomic data from the Rabi time in Sr1 are used only when scanning over Sr1 Ramsey fringes. Prior to and during each scan, the Rabi data measure frequency drift of the free-running LO (typically between 0 and 3 mHz s⁻¹), and allows us to apply LO drift compensation in a double-integrator control loop with attack time of approximately 100 clock cycles. (b) Results using Ramsey spectroscopy in Sr1, cointerrogated with Sr2 using the same LO. Left: excitation fractions with the LO frequency locked to the central Sr1 Ramsey fringe under three conditions: 300 ms Ramsey dark time with the APL to Sr2 disengaged (orange), 2 s dark time with the APL disengaged (red), and 2 s dark time with the APL engaged, the fringe width is measured to be 254(1) mHz for a 2 s dark time.

proposed [50] and demonstrated [51] for microwave atomic clocks. An initial 10 ms $\pi/2$ pulse drives the atomic ensemble into a coherent state on the equator of the Bloch sphere with $\langle S_x \rangle = N/2$. As in a normal Ramsey sequence, the atomic state is left to freely evolve during which time it accumulates a phase shift relative to the LO. In the small angle approximation the accumulated LO phase is proportional to $\langle S_{v} \rangle$, which is read out in the following procedure: a $\pi/2$ pulse is driven by the LO, the phase of which is stepped by 90° with respect to the initial pulse in order to map S_{y} to S_{z} . The ground state atom number S_{qq} is then read out via a QND measurement pulse with duration $t = 30 \ \mu s$. To read out the excited state atom number S_{ee} , a π pulse is driven with LO phase -90° relative to the initial pulse, before a second QND measurement is applied for $t = 30 \ \mu s$. Finally, the LO phase is stepped again to 90° and a final $\pi/2$ pulse is applied to return the collective atomic spin to the equator of the Bloch sphere. Based on the results of the two OND measurements, the LO phase is stepped to align the atomic spin to point along the x axis of the Bloch sphere. Repeating the free-evolution time and the QND measurement procedure several times in succession, a phase lock of the LO to the atomic transition can be maintained for several seconds-well beyond the coherence time of the free-running LO.

To characterize the improvement in LO phase noise, the Sr2-phase-stabilized light is used to interrogate Sr1, with results shown in Fig. 4(b). Atomic samples are prepared in parallel in both systems and probed synchronously using the same local oscillator. Sr2 is used to implement the APL while Sr1 performs standard Ramsey spectroscopy. To get a baseline measurement of the free-running laser phase noise, the Sr2 APL is first disengaged and Sr1 is operated as a clock with Ramsey spectroscopy dark time T = 300 ms. When we lock the frequency of the LO to the central Sr1 fringe, we observe noise in the in-lock excitation fraction corresponding to a standard deviation for the accumulated LO phase error of 290 mrad. Increasing Ramsey dark time to 2 s, but with the APL still disengaged, shows that the accumulated phase error is too large to operate the clock reliably. This is clear from the S-shaped histogram of the excitation noise in Fig. 4, indicating the phase error is well outside the $\pm \pi/2$ range that is needed to determine unambiguously the average frequency offset during the dark time. A final dataset is taken with the APL engaged on Sr2 during the 2 s Ramsey dark time in Sr1. Specifically, the APL consists of five repetitions of a 340 ms dark time followed by 60 ms QND phase measurement and correction. The phase corrections are applied onto an acoustooptic modulator (AOM) which corrects the LO light prior to it being split and sent to both clocks. Therefore, the average residual phase error accumulated during the APL can be characterized based on the excitation noise in Sr1 when locked to the clock transition, and was determined to be 240 mrad. Finally, a scan of the full Ramsey fringe in Sr1

shows no degradation of the 95% contrast and a Fourier limited linewidth of 254(1) mHz, corresponding to an oscillator Q factor of 1.7×10^{15} . This is within a factor of 3 of the finest scan resolution achieved using state-of-the-art LOs, but unlike earlier high-resolution scans [28,52,53] we observe no significant loss of contrast on the Ramsey fringes. To our knowledge this matches the narrowest spectroscopic feature to which any oscillator has yet been stabilized [54]. Extending the APL time further, either through longer free-evolution time or increased number of QND measurements, resulted in increased phase noise in Sr1.

V. CONCLUSION

We have demonstrated that the QND-based APL is an effective approach to improve the phase coherence of an ultrastable laser, making it a competitive alternative to other strategies for minimizing the technical noise of the LO [55]. Increasing the LO phase coherence directly impacts the frequency stability performance of the OLC, as it enables longer Ramsey dark time T, resulting in an increased Q factor and a steeper discriminant of the atomic excitation fraction against the LO frequency. The clearest impact of this is on the quantum-projection-noise- (QPN) induced fractional frequency instability, which for spectroscopy of N atoms with a signal contrast C and a cycle time of T_c is given by

$$\sigma_{\rm QPN}(\tau) = \frac{1}{\pi QC} \sqrt{\frac{T_c}{N\tau}},\tag{7}$$

where $\sigma(\tau)$ denotes the Allan deviation for averaging time τ in seconds. Specifically for Sr1, which operates with 5×10^3 atoms, the sevenfold increase in the *Q* factor achieved by extending the Ramsey probe time from 300 ms to 2 s, with corresponding cycle times 1.3 and 3 s respectively, reduces the QPN instability from $2.1 \times 10^{-17}/\sqrt{\tau}$ to $4.8 \times 10^{-18}/\sqrt{\tau}$.

Another important source of instability in OLCs is the Dick effect, caused by short-term LO frequency noise which is sampled by dead time (primarily cooling time) in the clock sequence. For Ramsey spectroscopy, in the limit of instantaneous $\pi/2$ pulses, the Dick-effect instability is given by [56]

$$\sigma_{\text{Dick}}(\tau) = \sqrt{\frac{1}{\tau} \sum_{k=1}^{\infty} S_y(k/T_c) \left[\frac{\sin(\pi kT/T_c)}{\pi kT/T_c}\right]^2}.$$
 (8)

Increasing the ratio of the Ramsey dark time to the cycle time helps to suppress this effect. Estimating the precise reduction in Dick-effect instability is complex, as it depends on the power spectral density $S_y(f)$ of the fractional frequency fluctuations of the LO at harmonics of the cycle frequency. Under the assumption that LO flicker noise, which we have directly measured to be 8×10^{-17} , is the dominant noise process, the estimated Dick-effect induced instability for a 2 s Ramsey dark time is $5 \times 10^{-17}/\sqrt{\tau}$ —a factor of 1.6 below what is expected for a 300 ms Ramsey dark time, leading to a reduction in measurement time by a factor of 2.5 to reach the same precision. In the future, an optical frequency comb could transfer the enhanced phased stability of the LO to other wavelengths in order to improve the performance of optical clocks based on different atomic species [57,58]. In particular, applying this technique to Yb⁺ or highly charged ion clocks which are limited by QPN, but exhibit a large sensitivity to changes in the fine structure constant, could facilitate improved tests of fundamental physics [59–61].

It is instructive to compare our QND-based method against alternative approaches to extend the coherence time of the LO laser. In one demonstration, the LO was prestabilized to an OLC with 50% duty cycle, and then used to interrogate a second OLC [28]. However, with this approach the attainable extension of probe time is limitedthere is still considerable dead time of several hundred milliseconds needed to cool atoms in the prestabilization OLC, during which the phase of the LO is going unmeasured. Another recent demonstration used a novel multipulse interrogation scheme in an OLC to achieve dynamical decoupling of the laser phase noise. The dynamical decoupling allows the OLC to coarsely track the laser phase for a continuous spectroscopy time much longer than the coherence of the LO [62]. However, the dynamical decoupling also reduces the mean sensitivity of the OLC to laser phase fluctuations, resulting in a higher sensitivity to quantum projection noise compared with the QND-based scheme presented in this work. Finally, another promising alternative would be to make use of recent advances combining strontium atoms and tweezer arrays [53,63]. Such platforms allow for repeated probing of the clock transition and detection, in some cases up to 15 times, without needing to reload the atoms. However, since these experiments rely on fluorescence detection, the phase coherence between the LO and the atoms is lost during detection. If repeated fluorescence readout in a tweezer array were used to implement a destructive form of the atom phase lock, the phase measurement errors (e.g., from quantum projection noise) would accumulate with each interrogation pulse. In contrast, the QND measurementbased approach preserves coherence after each measurement, resulting in phase noise in earlier measurements being corrected for by subsequent measurements.

Considering alternative applications of the APL scheme, we speculate that it could help to enable high-Q spectroscopy in environments where the ultimate performance of cavity stabilized lasers cannot be reached, for instance in field deployed systems. At the same time, the QND measurement scheme underpinning the APL also opens the door to other configurations of quantum-enhanced optical

atomic clocks. Going forward, it will be instructive to characterize the QND measurement scheme in more detail, for example, by using colder atomic samples in a bettercontrolled motional state. Relative to the data presented in Fig. 2, we observe that the QND contrast decay rate can be reduced by a factor of approximately 2 by adjusting the MZI setup in Fig. 1 so that the two electro-optic modulators (EOMs) generating the stronger probe sidebands at $\Omega/2 \pm$ 125 MHz are placed on different arms of the MZI-this eliminates spurious second-order frequency components near atomic resonance created at the difference frequency between those EOMs, thereby mitigating a source of excess scattered photons. If the OND readout noise and measurement backaction can be controlled close to their shot noise limits, our quantum nondemolition measurement apparatus could be used to generate squeezed states with reduced QPN, offering a route toward OLC comparison with unprecedented frequency precision. Finally, the ability to engineer squeezing in Sr could also have implications beyond precision timekeeping, for example, by improving the performance of Sr atom interferometers [64] toward the sensitivity necessary to observe gravitational waves [65].

ACKNOWLEDGMENTS

This work was financially supported by the UK Department for Business, Energy and Industrial Strategy as part of the National Measurement System Programme, and by the European Metrology Programme for Innovation and Research (EMPIR) Project No. 17FUN03-USOQS. This project has received funding from the EMPIR programme cofinanced by the Participating States and from the European Union's Horizon 2020 research and innovation programme. A. V. acknowledges funding from the Engineering and Physical Sciences Research Council (EPSRC UK) through the Controlled Quantum Dynamics Centre for Doctoral Training (EP/L016524/1) for the core duration of this work. We thank Helen Margolis, Alissa Silva, and Jake Paterson for operating the optical frequency comb, and Rachel Godun and Rich Hendricks for careful reading of the manuscript.

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